

**Nonlocality of the isobar propagation and the effective  $\Delta$ -nucleus spin-orbit interaction**

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The imaginary central and spin-orbit components of the  $\Delta$ -nucleus optical potential are investigated in a  $\pi$ -exchange model. It is found that the effective spin-orbit interaction of the isobar reflects large nonlocalities from true pion absorption and from the quasielastic  $\Delta$  decay. The parameters for the absorption and the spin-orbit strength are in qualitative agreement with recent phenomenological results.

[ NUCLEAR SCATTERING  $\Delta$  isobar propagation; relation of nonlocality ]  
 and spin-orbit interaction.

The formulation of the isobar doorway model<sup>1-7</sup> has provided new insight into the dynamics of nuclear reactions at intermediate energies. In particular it turned out that in view of the selective sensitivity of the  $\Delta$ -particle  $N$ -hole ( $\Delta\bar{N}$ ) states on the nuclear medium, higher order corrections—beyond that of pion multiple rescattering—are crucial for a quantitative understanding. Unfortunately, a microscopic cal-

ulation of such medium corrections—among which true  $\pi$ -absorption and reflection terms are the most important ones<sup>3</sup>—requires rather drastic assumptions about the elementary input. Alternatively Hirata *et al.* and recently Horikawa *et al.* parametrized the higher order corrections in terms of a local one-body operator with a central<sup>2</sup> and an effective  $\Delta$ -nucleus spin-orbit interaction<sup>7</sup>

$$V_{\Delta}(r; \omega) = \left\{ [V_c(\omega) + iW_c(\omega)] - [V_{ls}(\omega) + iW_{ls}(\omega)] \bar{\Gamma} \bar{\Sigma} \frac{r_0^2}{r} \frac{d}{dr} \right\} \rho(r) / \rho_0 \quad (1)$$

as a function of the scattering energy  $\omega$  and the nuclear density  $\rho(r)$  ( $r_0$  is a scale parameter, taken to be 1 fm). For an energy-independent spin-orbit term  $W_{ls}(\omega)$  a best fit of elastic  $\pi$ -scattering data on <sup>12</sup>C was obtained with (compare Fig. 3)

$$W_c(\omega) \approx -40 \text{ MeV} \quad (2)$$

together with

$$W_{ls}(\omega) \approx -W_c(\omega) / 10 \quad (3)$$

Though the ansatz in Eq. (1) is fairly successful in actual calculations, its shortcomings are obvious: Incorporating different medium corrections in such a simple parametrization necessarily prevents their detailed and systematic investigation. In addition, an interpretation of the various effective coupling constants, obtained from a fit of Eq. (1) to experimental data, is not unique, as the resulting parameters have to reflect the shortcomings of the parametrization itself [for example, they artificially have to mock-up nonlocal effects in the central part of the isobar-nucleus optical potential, absent in the ansatz in Eq. (1)].

In a microscopic approach to the parametrization from Eq. (1) we concentrate in the following on the absorptive part of  $V_{\Delta}(r; \omega)$ : for this piece a diagrammatic expansion is promising due to the small number of inelastic channels (in addition, the real part of the  $\Delta$ -nucleus potential receives large contri-

butions from  $\sigma$  and  $\omega$  exchange and requires a much more detailed model than sketched below).

In the same spirit we evaluate the two leading diagrams from Fig. 1 for  $\pi$  exchange only, mocking-up the influence of heavy mesons, such as the  $\rho$  meson, by a relatively small  $\pi$  cutoff with  $\Lambda_{\pi} \sim 800 \text{ MeV}$ .<sup>8</sup> With static  $\pi NN$  and  $\pi N\Delta$  vertex functions we then obtain from old-fashioned perturbation theory for the

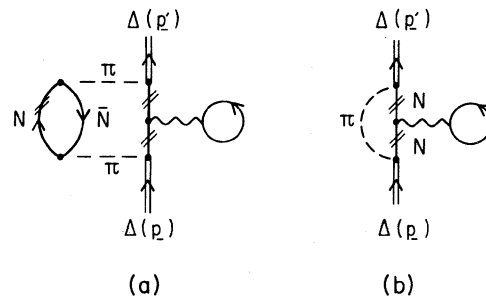


FIG. 1. Leading diagrammatic contributions to the imaginary isobar self-energy (to first order in the nuclear density) through the coupling to the 2p-1h continuum due to true  $\pi$  absorption ( $\pi, 2N$ ) (a) and to the quasielastic channel (b). Above the wiggly lines represent the shell model potential of the nucleon; the double lines indicate Pauli blocking for the nucleon.

2 $\pi$ -exchange diagram in Fig. 1(a)

$$V_{\pi\pi}(\vec{p}, \vec{p}'; \omega) = \left(\frac{f_\pi}{m_\pi}\right)^2 \left(\frac{f_\pi^*}{m_\pi}\right)^2 \frac{1}{(2\pi)^3} \int \frac{d\vec{k} N(\vec{k}; \vec{p}, \vec{p}')}{\omega_\pi(\vec{k} + \vec{p}) \omega_\pi(\vec{k} + \vec{p}')} \frac{F_\pi^2(\vec{k} + \vec{p}') F_\pi^2(\vec{k} + \vec{p}) F_\Delta(k_0)}{2T_N(k) - \omega + i\epsilon}$$

$$\times \frac{1}{2} \left\{ \frac{1}{\omega_\pi(\vec{k} + \vec{p}') + T_N(k)} + \frac{1}{\omega_\pi(\vec{k} + \vec{p}') + T_N(k) - \omega + i\epsilon} \right\}$$

$$\times \frac{1}{2} \left\{ \frac{1}{\omega_\pi(\vec{k} + \vec{p}) + T_N(k)} + \frac{1}{\omega_\pi(\vec{k} + \vec{p}) + T_N(k) - \omega + i\epsilon} \right\} \quad (4)$$

with the numerator

$$N(\vec{k}; \vec{p}, \vec{p}') = \vec{S}_1^\dagger(\vec{k} + \vec{p}') \vec{\sigma}_2(\vec{k} + \vec{p}') \vec{S}_1(\vec{k} + \vec{p}) \vec{\sigma}_2(\vec{k} + \vec{p}) \vec{T}_1^\dagger \vec{\tau}_2 \vec{T}_1 \vec{\tau}_2. \quad (5)$$

Above  $f_\pi$  and  $f_\pi^* = 2f_\pi$  (Ref. 9) denote the  $\pi NN$  and  $\pi N\Delta$  coupling constants;  $\omega_\pi(q) = (q^2 + m_\pi^2)^{1/2}$  is the pion energy and  $T_N(k)$  the kinetic energy of a nucleon in the intermediate state (corrections on the level of single particle energies are dropped). The form factors  $F_{\pi(\Delta)}(q)$  (Ref. 9) in Eq. (4) include off shell corrections both for the virtual pions (we use the same cutoff mass  $\Lambda_\pi$  at the the  $\pi NN$  and the  $\pi N\Delta$  vertices) and for the  $\Delta$ -isobar<sup>10,11</sup> ( $k_0$  is the on shell pion momentum in the  $\pi N$  c.m. system for the scattering energy  $\omega$  with  $k_0 = k_R \sim 230$  MeV/c at resonance).

The kinematics of the two-nucleon process allows a rather natural separation of the nonlocal potential  $V_{\pi\pi}(\vec{p}, \vec{p}'; \omega)$  of Eq. (4) into a local piece and the leading nonlocal correction: At resonance the momentum of the nucleons emitted satisfies

$$k_N \approx \sqrt{M\omega} \approx 500 \text{ MeV}/c \gg p, p' \quad (6)$$

for typical momenta  $p$  and  $p'$  of the isobar. By dropping then all the  $\vec{p}$  and  $\vec{p}'$  dependence relative to  $\vec{k}_N$  except for the invariant  $\vec{S}_1^\dagger \vec{p}' \vec{S}_1 \vec{p}$  the only dependence of  $V_{\pi\pi}(\vec{p}, \vec{p}'; \omega)$  on the isobar momentum is

$$W_c(\omega) = -\frac{1}{48\pi} \left(\frac{f_\pi}{m_\pi}\right)^2 \left(\frac{f_\pi^*}{m_\pi}\right)^2 \frac{Mk_N^5}{\omega_\pi^2(k_N)} F_\pi^4(k_N) F_\Delta(k_0) \rho_0 \left[ \frac{1}{\omega_\pi(k_N) + E_N(k_N) - E_\Delta} + \frac{1}{\omega_\pi(k_N) + E_N(k_N) - E_N} \right]^2 \quad (10)$$

and

$$W_b^1(\omega) = -W_c(\omega)/2k_N^2. \quad (11)$$

(Above  $\rho_0$  denotes the nuclear density at the origin.)

The two leading corrections to Fig. 1(a) come from the influence of the nuclear shell model potential and from the Pauli blocking.<sup>2,3,12</sup> While the first contribution increases  $W_b(\omega)$  by

$$\Delta W_b^1(\omega) \cong \left[ \frac{1}{6} \frac{M W_c(\omega)}{k_N^2} \right] V_b^N \quad (12)$$

[ $V_b^N \cong 18$  MeV (Ref. 13) is the coefficient for the

contained in the spin-orbit term of the numerator, which then reads, after summing over the intermediate  $NN$  state, as

$$N(\vec{k}; \vec{p}, \vec{p}') = \frac{k^2}{3} \left[ k^2 + \frac{i}{2} \underline{\Sigma}(\vec{p}' \times \vec{p}) \right], \quad (7)$$

where the transition matrix  $\underline{\Sigma}$  is defined by  $\langle \frac{3}{2} \parallel \underline{\Sigma} \parallel \frac{3}{2} \rangle = 2\sqrt{15}$ .

In going over to coordinate space we introduce the spin-orbit operator for the isobar by

$$i \underline{\Sigma}(\vec{p}' \times \vec{p}) \rightarrow \vec{T} \underline{\Sigma} \frac{1}{r} \frac{d}{dr}. \quad (8)$$

As the remaining part of the Box diagram is independent of  $\vec{p}$  and  $\vec{p}'$  its Fourier transform yields schematically

$$V_{\pi\pi}(\vec{r} - \vec{r}_N; \omega) \sim \delta(\vec{r} - \vec{r}_N). \quad (9)$$

By folding in the nuclear density we recover the form of Eq. (1) for the  $\Delta$ -nucleus potential with the imaginary central and spin-orbit part given, respectively, as

spin-orbit interaction of the nucleon], Pauli blocking reduces the total spin-orbit strength in  $^{12}\text{C}$  by a factor of

$$P(k_N) \cong 1 - \frac{\sqrt{2}}{18} \frac{7 + 4a^2(k_N - \sqrt{2}/a)^2}{ak_N} \times \exp \left[ -a^2 \left( k_N - \frac{\sqrt{2}}{a} \right)^2 \right] \quad (13)$$

for a  $\Delta$  isobar with an average momentum  $(\langle p^2 \rangle)^{1/2} \cong (\sqrt{2}/a)$  [ $a = 1.69$  fm (Ref. 7)]. The effect of both corrections is very small due to their mutual cancellation.

Similarly, we obtain for the spin-orbit contribution from Fig. 1(b) [the notation is the same as for Fig. 1(a)]

$$W_{ls}^2(r; \omega) = V_{ls}^N \frac{r_0^2}{r} \frac{d}{dr} \frac{\rho(r)}{\rho_0} \left( \frac{f_\pi^*}{m_\pi} \right)^2 (\bar{T}^\dagger \bar{T}) F_\Delta(k_0) \frac{1}{(2\pi)^3} \int \frac{d\vec{k} F_\pi^2(k)}{2\omega_\pi(k)} \frac{\bar{S}^\dagger \bar{k} \bar{\sigma} \bar{T} \bar{S} \bar{k}}{[\omega_\pi(k) + E_N(k) - E_\Delta + i\epsilon]^2} . \quad (14)$$

Evaluating the integral with standard techniques we obtain with

$$\bar{S}^\dagger \bar{k} \bar{\sigma} \bar{T} \bar{S} \bar{k} = \frac{k^2}{9} \bar{\Sigma} \bar{T} \quad (15)$$

the spin-orbit coefficient

$$W_{ls}^2(\omega) = \frac{\Gamma_\Delta(\omega)}{2} \left( \frac{\omega}{k_0^2} - \frac{1}{3} \frac{1}{M + \omega} \right) V_{ls}^N P(k_0) , \quad (16)$$

where  $\Gamma_\Delta(\omega)$  is just the width of the  $\Delta$  isobar at the scattering energy  $\omega$ ,<sup>14</sup> while  $P(k_0)$  again accounts for Pauli blocking of the intermediate nucleon.

Our main findings for <sup>12</sup>C are presented in Figs. 2 and 3. In Fig. 2 the energy dependence of  $W_{ls}^1(\omega)$  and  $W_{ls}^2(\omega)$  is shown for representative values of  $\Lambda_\pi$  and  $\Lambda_\Delta$ . Characteristically, the two coefficients show an opposite trend with increasing  $\omega$  as expected from their gross structure

$$W_{ls}^1(\omega) \propto 1/\omega; \quad W_{ls}^2(\omega) \propto \omega^{3/2} . \quad (17)$$

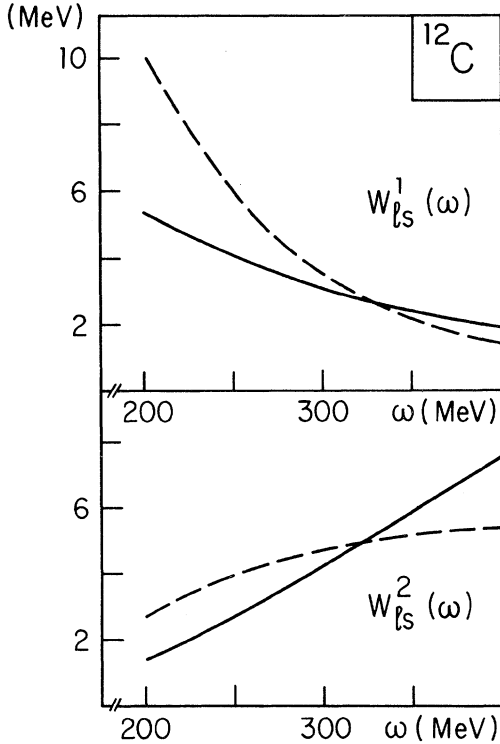


FIG. 2. Energy dependence of the coefficients  $W_{ls}^1(\omega)$  and  $W_{ls}^2(\omega)$  (without Pauli corrections) for  $\Lambda_\pi = 800$  MeV and two different cutoff masses  $\Lambda_\Delta = \infty$  (full lines) and  $\Lambda_\Delta = 200$  MeV (dashed lines), respectively.

For quantitative details we consider for  $\Lambda_\pi = 800$  MeV (Ref. 8) two extreme choices with  $\Lambda_\Delta = \infty$  (no cutoff) and  $\Lambda_\Delta = 200$  MeV (Ref. 11) for the off shell continuation of the  $\Delta$  isobar. The influence of  $\Lambda_\Delta$  is significant; presently, its uncertainty just reflects current problems in defining the  $\Delta$  isobar microscopically, especially off the resonance (for different philosophies compare, for example, Refs. 15–17).

In Fig. 3 we compare our results with the findings by Horikawa *et al.* For  $\Lambda_\pi \sim 800$  MeV without an additional  $\Delta$  cutoff we qualitatively reproduce the  $W_c(\omega)$  from Ref. 7. As the same model fits the total  $\pi$ -absorption cross section  $\pi d \rightarrow NN$  (Ref. 8) we conclude from our qualitative agreement that the phenomenological quantity  $W_c(\omega)$  in Ref. 7 receives its dominant contribution from the true  $\pi$ -absorption process ( $\pi, NN$ ). For  $\Lambda_\Delta = 200$  MeV the agreement is much worse as the resulting energy dependence of  $W_c(\omega)$  is too steep. The same situation persists for

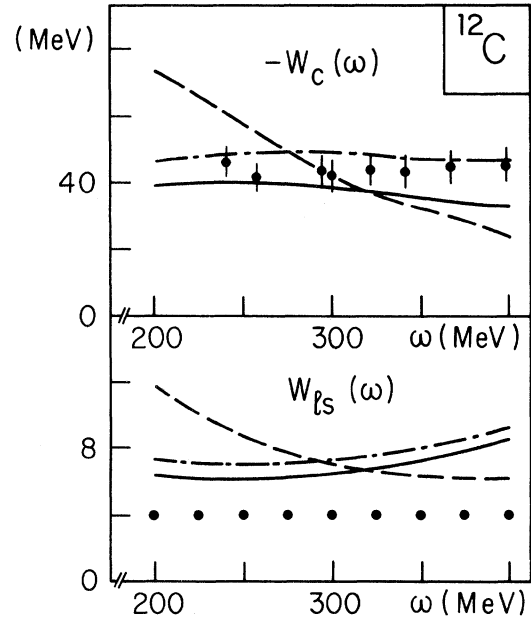


FIG. 3. Energy dependence of true  $\pi$  absorption  $W_c(\omega)$  [Eq. (11)] and the summed strength  $W_{ls}(\omega) = W_{ls}^1(\omega) + W_{ls}^2(\omega)$  of the spin-orbit interaction of the  $\Delta$  isobar. Shown are the results for  $\Lambda_\pi = 800$  and 900 MeV without  $\Delta$  cutoff (full and dashed-dotted lines) and for  $\Lambda_\pi = 800$  MeV together with  $\Lambda_\Delta = 200$  MeV (dashed line), as well as the corresponding results from Horikawa *et al.* (Ref. 7) (full dots).

the total spin-orbit strength  $W_b(\omega) = W_b^1(\omega) + W_b^2(\omega)$ : only without a  $\Delta$  cutoff we find  $W_b(\omega)$  approximately constant as in the fit from Ref. 7. For the discrepancy in the absolute magnitude by a factor 1.5–2 there are two obvious interpretations: on the one side we presumably overestimate  $W_b(\omega)$ , as our model incorporates neither  $\rho$ -exchange nor short range correlations [both mechanisms should cut down  $W_b(\omega)$ ]; on the other side,  $W_b(\omega)$  from Ref. 7 might indeed underestimate the complex spin-orbit interaction of a  $\Delta$  isobar, as  $\pi$  scattering in the isobar-hole model is not yet understood on such a quantitative level.

Summarizing, we can account for some aspects of the parametrization of Ref. 7 in our simple microscopic model. We find that for a local parametrization of the central part of the  $\Delta$ -nucleus optical potential highly nonlocal medium corrections—dominated by true pion absorption and reflection—have to be mocked-up by a large spin-orbit interaction of the isobar. This explains why isobar doorway calculations, which keep the nonlocalities in the higher order corrections, are similarly successful without a strong spin-orbit term.<sup>3,18</sup> Furthermore, the result indicates that only within a more detailed

microscopic framework a comparison between the phenomenological spin-orbit interaction of the isobar from Ref. 7 and with the quark model,<sup>19</sup> for example, is meaningful.

It is clear that on a quantitative level the diagrammatic approach has its own problems. The sensitivity of the result on the cutoff masses is an unpleasant feature (though the same parameters already enter into a calculation of the  $\Delta N$  interaction in first order); more serious are the difficulties in developing on the same basis a quantitative picture for the real parts of the  $\Delta$ -nucleus potential, as it is not fully clear how well a diagrammatic expansion converges. For a conclusive answer further investigations have to be awaited.

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